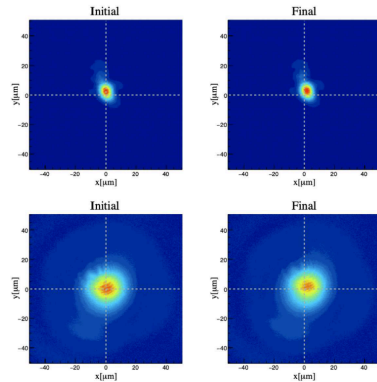
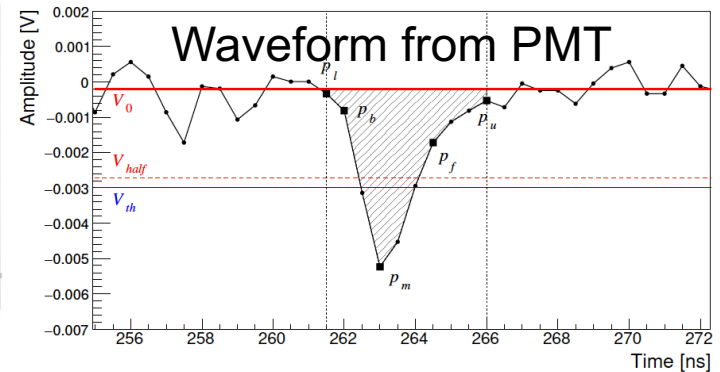
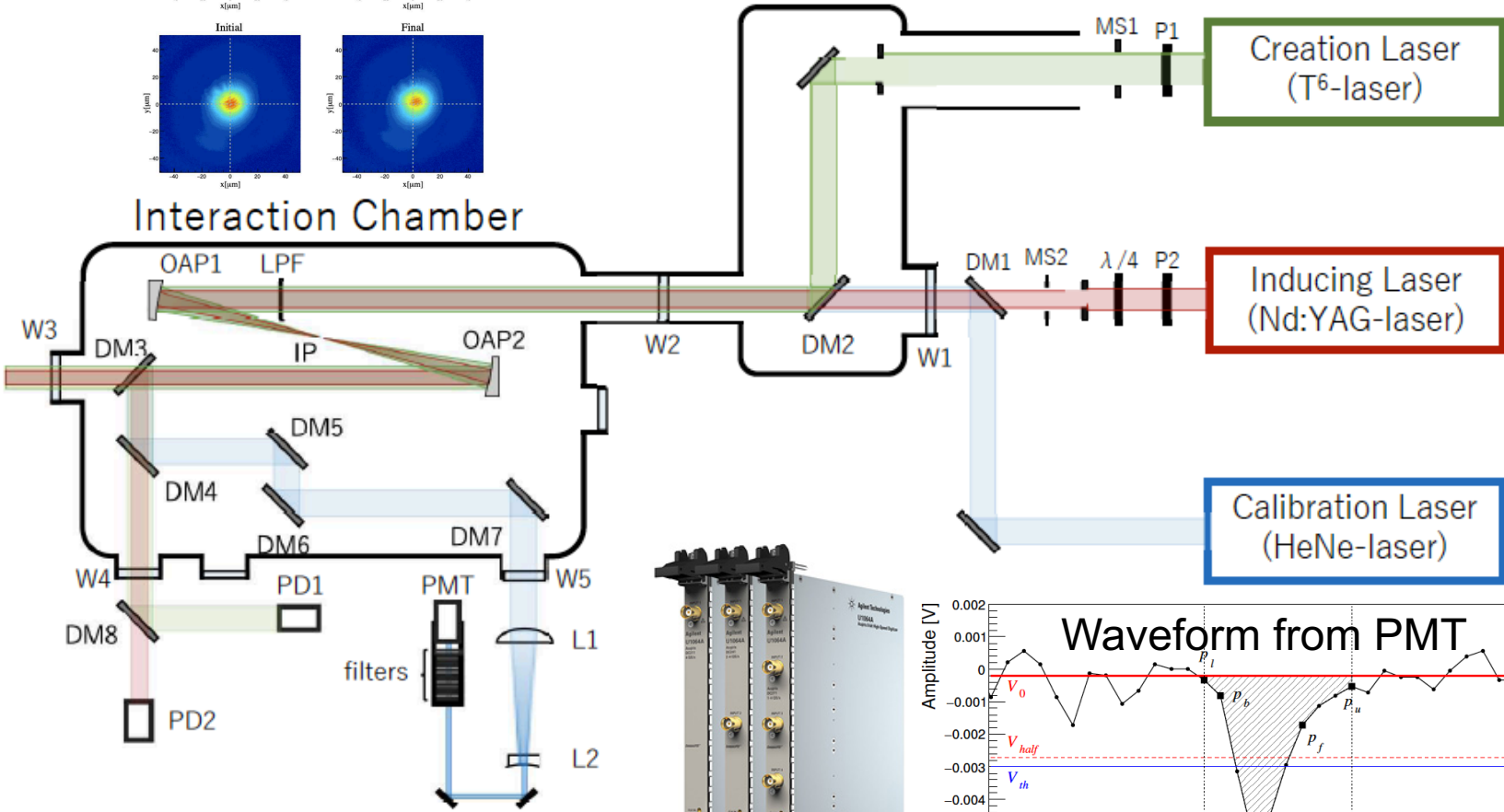


Searching setup



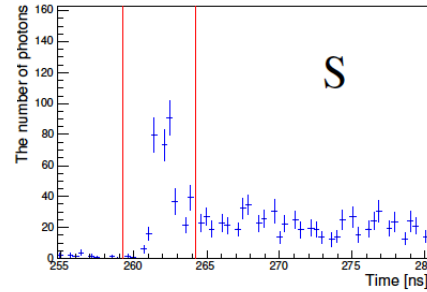
Transport Chamber



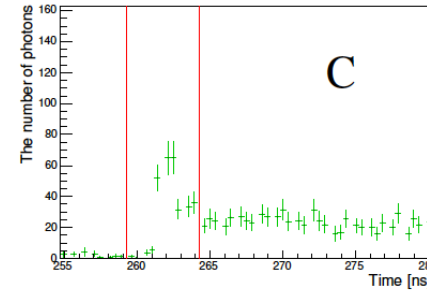
arXiv: 2105.01224
The SAPHIRES collaboration

Pressure dependence of atomic four-wave mixing signals

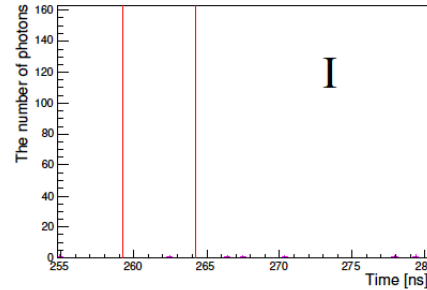
Creation +
Inducing lasers



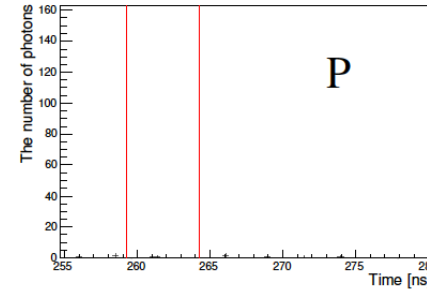
Creation laser



Inducing lasers

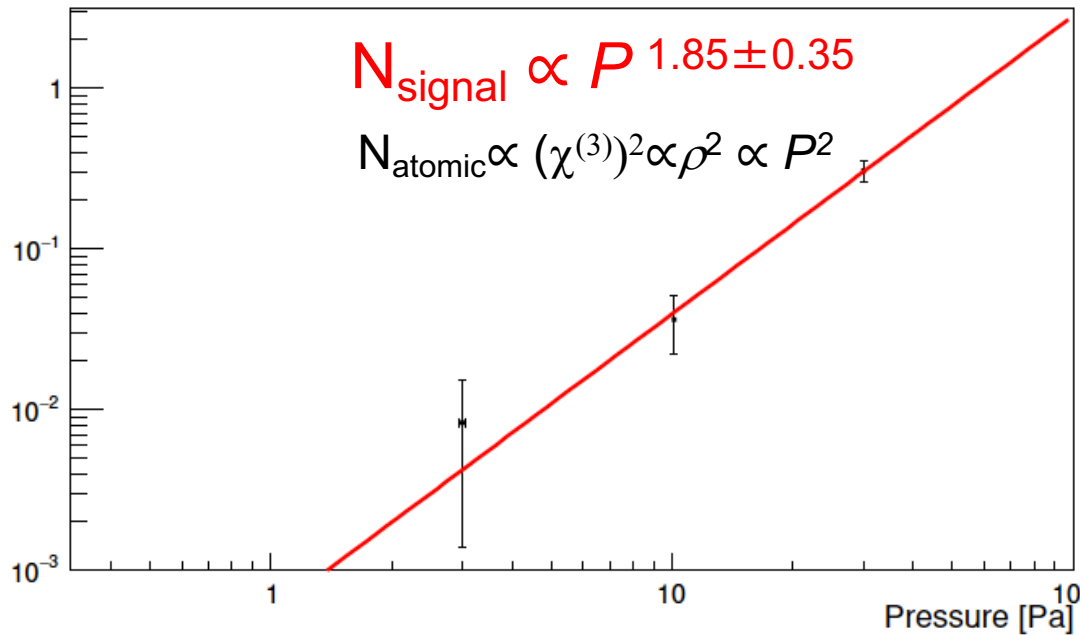


Pedestal



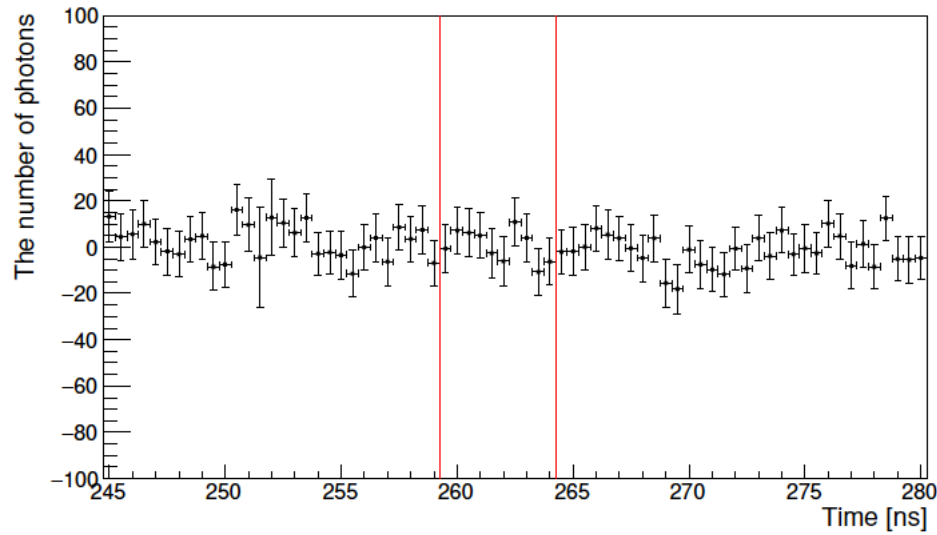
$$N_{\text{atomic}} = (S-P) - (C-P) - (I-P)$$

The number of photons per shot

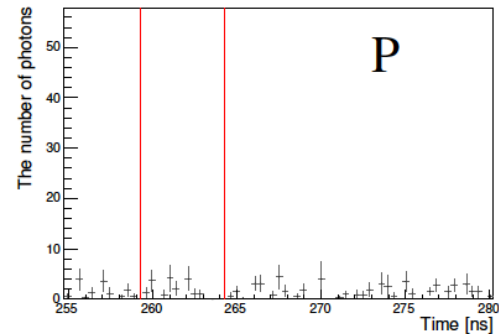
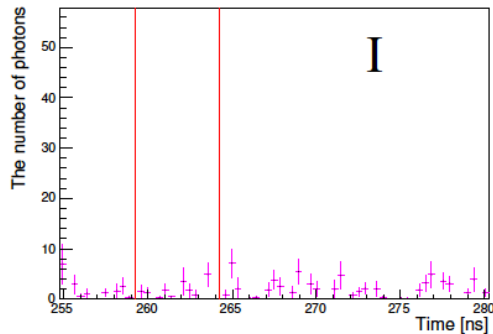
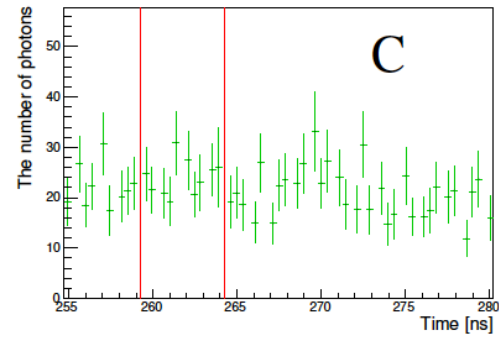
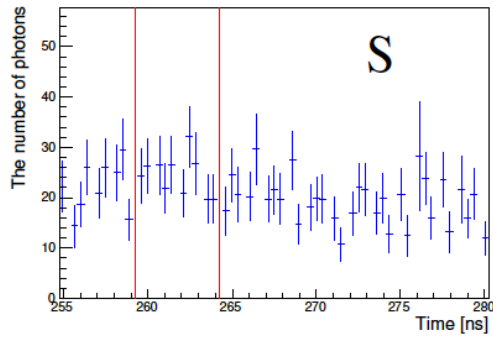


Search result at 2.6×10^{-5} Pa

$$N_{\text{signal}} = (S-P) - (C-P) - (I-P)$$



$$N_{\text{signal}} = 4.9 \pm 22.8(\text{stat.}) \pm 22.8(\text{sys.I}) \pm 3.8(\text{sys.II}) \pm 3.7(\text{sys.III})$$

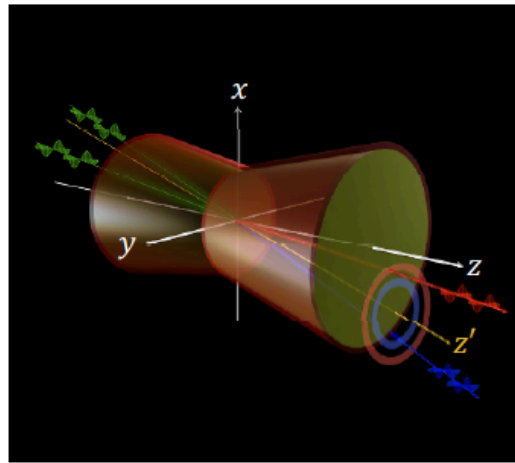
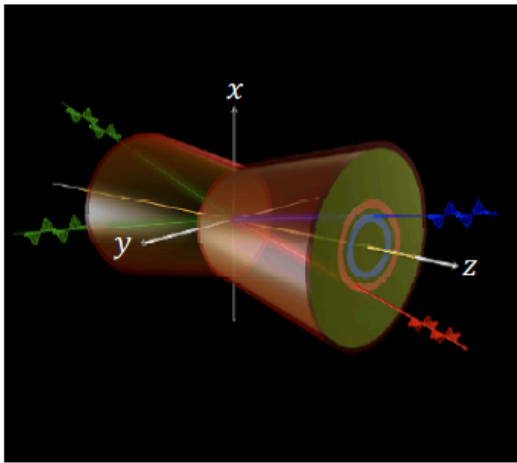
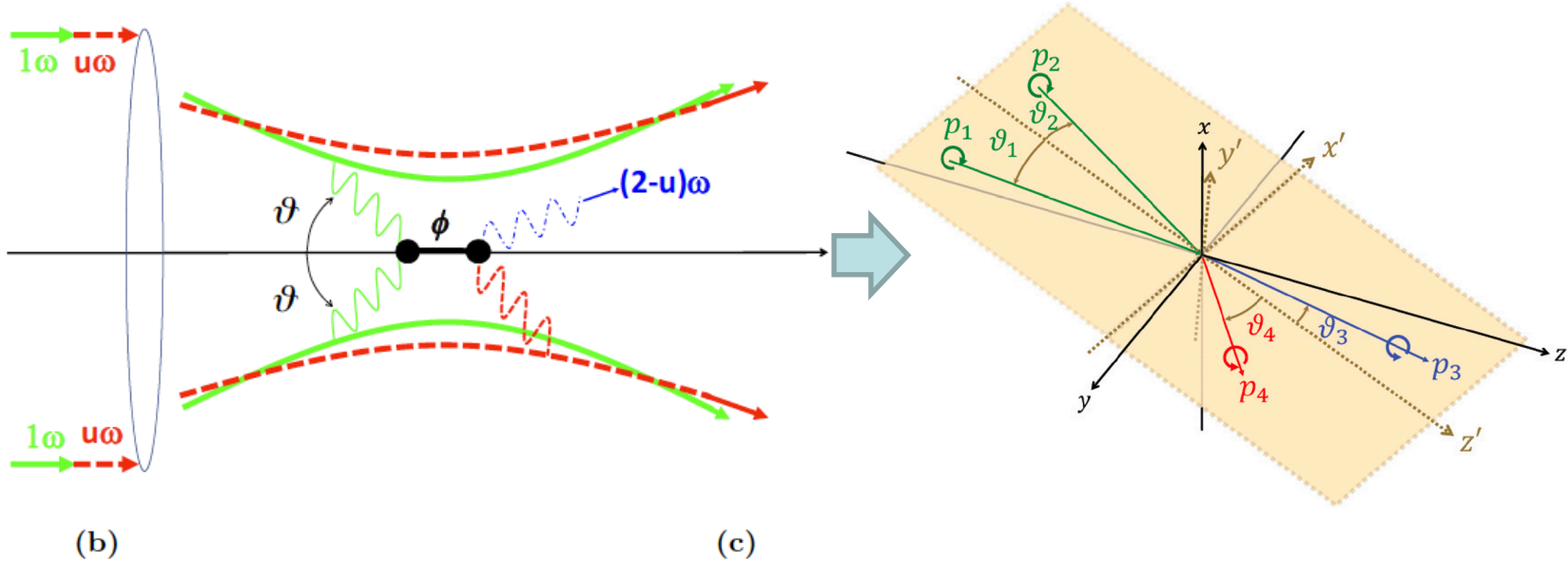


Extension to asymmetric-incident and non-coaxial collisions from symmetric geometry

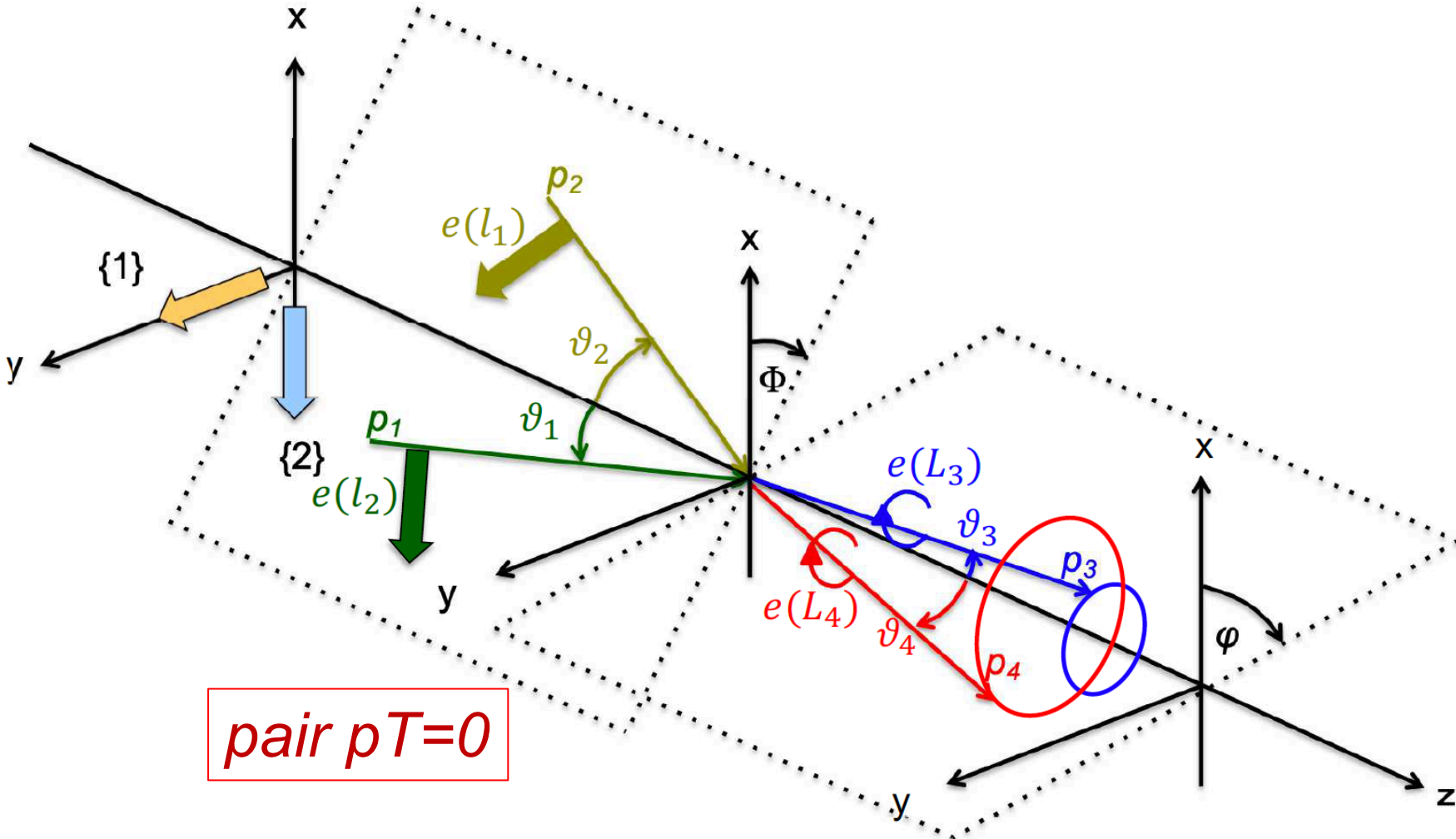
KH and Y. Kirita,

Stimulated radar collider for probing gravitationally weak coupling pseudo Nambu-Goldstone bosons, JHEP 09 95 (2020)

(a)



Linear polarization (creation) + Circular polarization (inducing)



Exclusion region

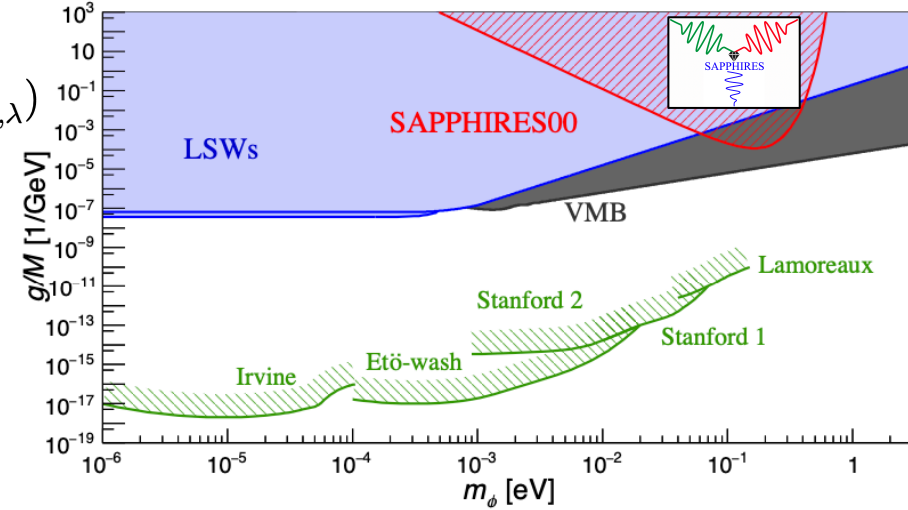
$$\mathcal{M}_S = \frac{1}{4} \left(\frac{g}{M} \right)^2 \frac{\text{Vertex Factors}}{m^2 - (p_1 + p_2)^2}$$

$$F^{\mu\nu} \equiv (-i) \int \frac{d^3\mathbf{p}}{(2\pi)^3 2p^0} \sum_{\lambda=1,2} (P^{\mu\nu} e^{-ipx} a_{\mathbf{p},\lambda} + \hat{P}^{\mu\nu} e^{ipx} a_{\mathbf{p},\lambda}^\dagger)$$

$$P^{\mu\nu} \equiv p^\mu e^\nu(p, \lambda) - e^\mu(p, \lambda) p^\nu,$$

$$\hat{P}^{\mu\nu} \equiv e^{*\mu}(p, \lambda) p^\nu - p^\mu e^{*\nu}(p, \lambda)$$

$$\text{Scalar-type : } (P_1(l_1) P_2(l_2)) (\hat{P}_3(L) \hat{P}_4(L))$$



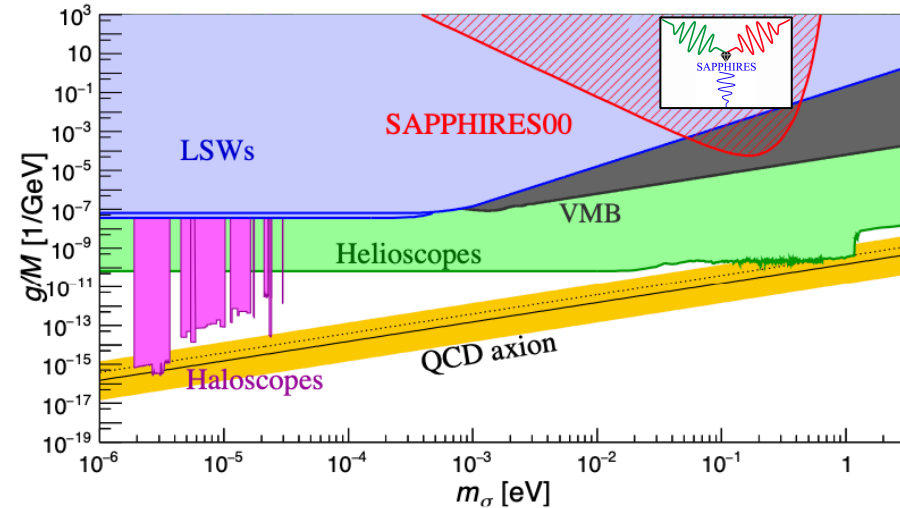
arXiv:2105.01224 (accepted by JHEP 2021)

$$\tilde{F}^{\mu\nu} \equiv \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} = (-i) \int \frac{d^3\mathbf{p}}{(2\pi)^3 2p^0} \sum_{\lambda=1,2} (\tilde{P}^{\mu\nu} e^{-ipx} a_{\mathbf{p},\lambda} + \tilde{\hat{P}}^{\mu\nu} e^{ipx} a_{\mathbf{p},\lambda}^\dagger)$$

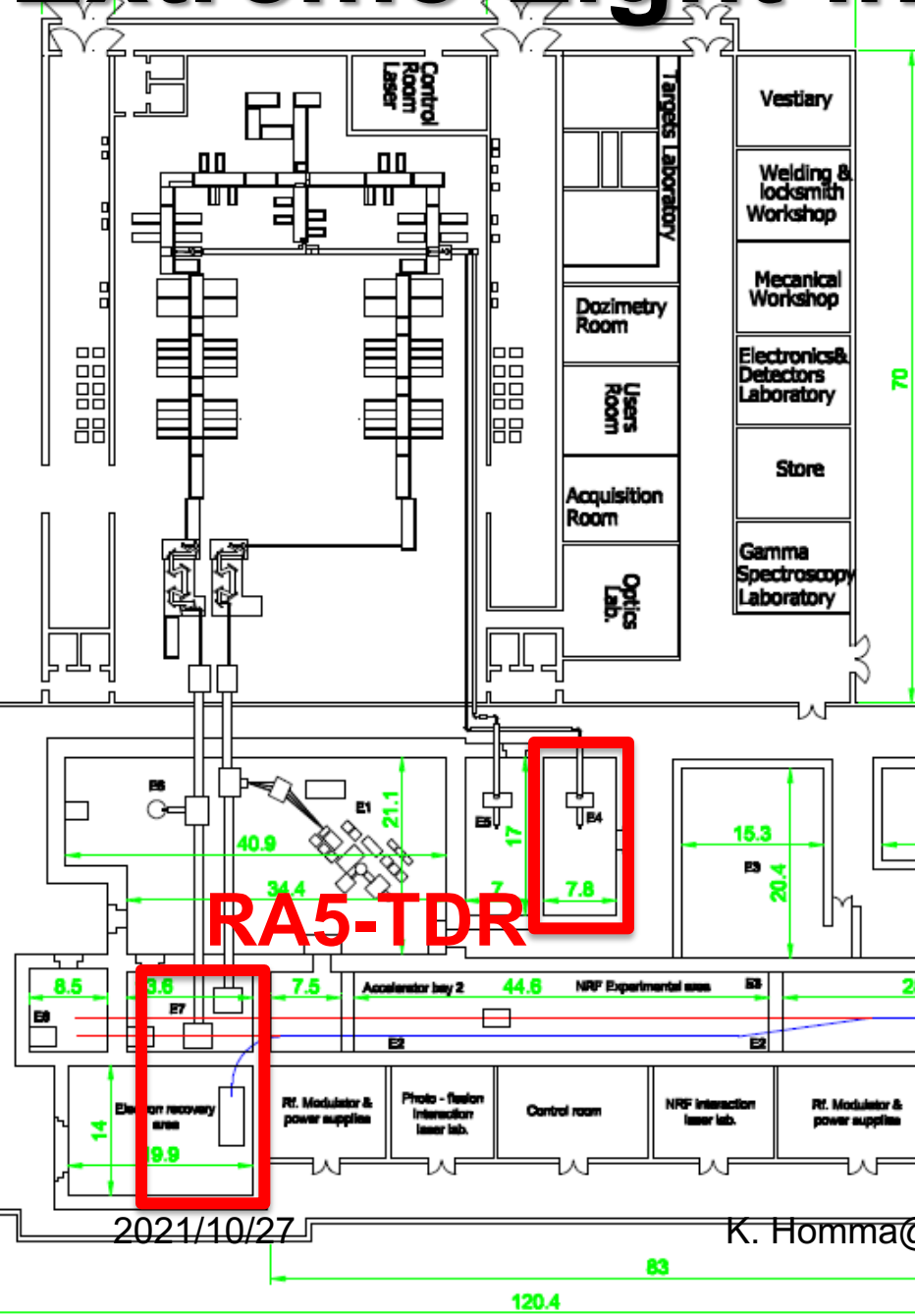
$$\tilde{P}^{\mu\nu} \equiv \epsilon^{\mu\nu\alpha\beta} (p_\alpha e_\beta(p, \lambda) - e_\alpha(p, \lambda) p_\beta)$$

$$\tilde{\hat{P}}^{\mu\nu} \equiv \epsilon^{\mu\nu\alpha\beta} (p_\alpha e_\beta^*(p, \lambda) - e_\alpha^*(p, \lambda) p_\beta)$$

$$\text{Pseudoscalar-type : } (P_1(l_1) \tilde{P}_2(l_2)) (\hat{P}_3(L) \tilde{\hat{P}}_4(L))$$



Extreme-Light-Infrastructure (ELI)



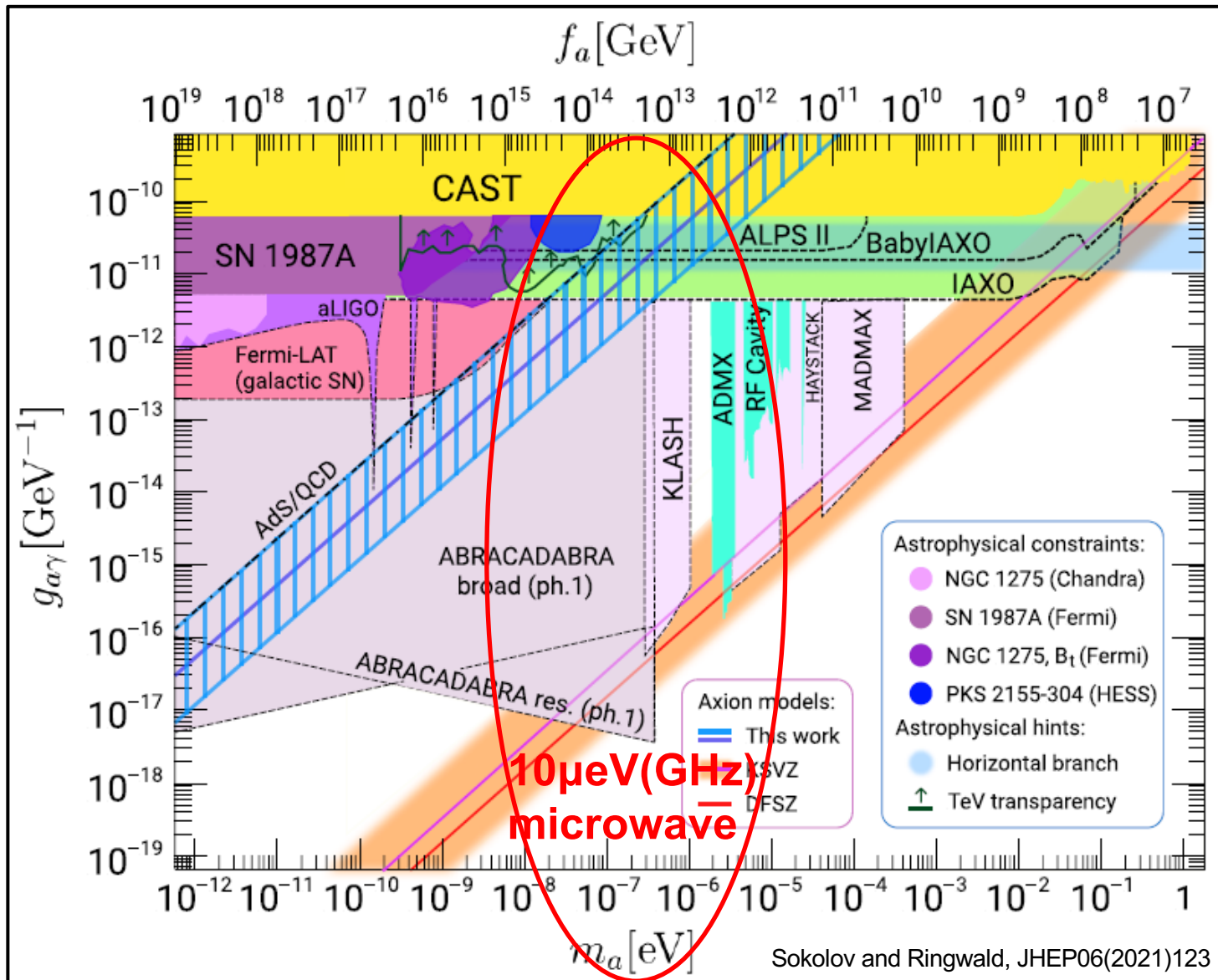
ELI-NP facility

- 2 x 10PW**
- 2 x 1 PW**
- 2 x 0.1 PW**

0.2-19.5 MeV gamma beam produced by ~700 MeV e- + laser

K. Homma@IITBWorkshop

Target mass-coupling domains



Dream of dilaton detection

Progress of Theoretical Physics, Vol. 126, No. 3, September 2011

An Approach toward the Laboratory Search for the Scalar Field as a Candidate of Dark Energy

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(Received March 10, 2011; Revised July 18, 2011)

The observed accelerating universe indicates the presence of Dark Energy which is probably interpreted in terms of an extremely light gravitational scalar field. We suggest a way to probe this scalar field which contributes to optical light-by-light scattering through the resonance in the quasi-parallel collision geometry. As we find, the frequency-shifted photons with the specifically chosen polarization state can be a distinct signature of the scalar-field-exchange process in spite of the extremely narrow width due to the gravitationally weak coupling to photons. Main emphasis will be placed in formulating a prototype theoretical approach, then showing how the weak signals from the gravitational coupling are enhanced by other non-gravitational effects at work in laser experiments.

dilaton as a candidate of dark energy

Conformal (dilatation) transform

$$g_{\mu\nu} \rightarrow g_{*\mu\nu} = \Omega^2(x)g_{\mu\nu}$$

Jordan Frame (JF) $\xi\phi^2 = (8\pi G_{eff})^{-1}$ $\phi = \hat{\xi}^{-1/2}\Omega$, with $\Omega = \exp(\hat{\zeta}\sigma)$
 $\hat{\xi} = \xi M_P^{-2}$, $\hat{\zeta} = \zeta M_P^{-1}$

$$\sqrt{-g} \left(\frac{1}{2}\xi\phi^2 R - \frac{1}{2}\epsilon g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi + L_{\text{matter}} - \Lambda \right)$$

Einstein Frame (EF)

Decaying $\Lambda \propto 1 / t^2$

$$\sqrt{-g_*} \left(\frac{1}{2}R_* - \frac{1}{2}g_*^{\mu\nu} \partial_\mu \sigma \partial_\nu \sigma + L_{*\text{matter}} - \Lambda \exp(-4\hat{\zeta}\sigma) \right)$$

$t=10^{60}$ in Planckian units

$\Lambda=10^{-120}$, $a(t)\propto t^{1/2}$

Coupling to Higgs field to realize constant mass

JF

EF

$$-\mathcal{L}_H = \sqrt{-g} \left(\frac{1}{2}h\phi^2\Phi^2 + \frac{\lambda}{4!}\Phi^4 \right) = \sqrt{-g_*}\Omega^{D-4} \left(\frac{1}{2}\tilde{m}^2\Phi_*^2 + \frac{\lambda}{4!}\Phi_*^4 \right)$$

$$\tilde{m}^2 = h\hat{\xi}^{-1}$$

pseudo-dilaton mass and coupling

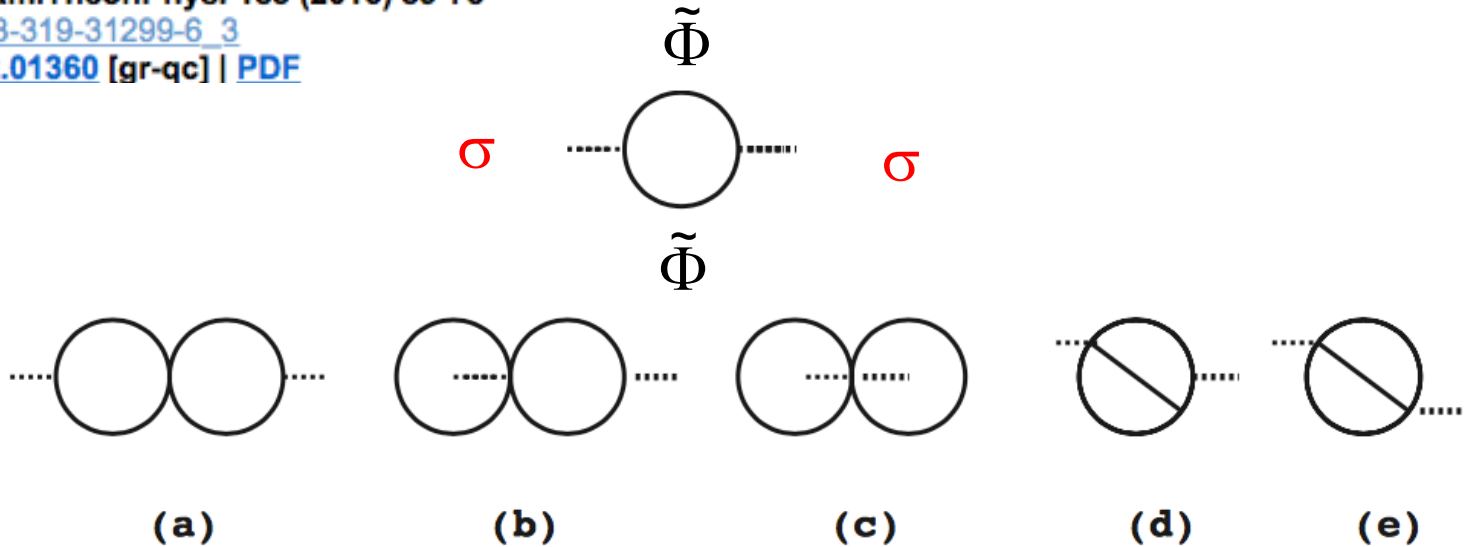
. A new estimate of the mass of the gravitational scalar field for Dark Energy

Yasunori Fujii (Waseda U., RISE). Dec 4, 2015. 17 pp.

Published in *Fundam.Theor.Phys.* 183 (2016) 59-75

DOI: [10.1007/978-3-319-31299-6_3](https://doi.org/10.1007/978-3-319-31299-6_3)

e-Print: [arXiv:1512.01360](https://arxiv.org/abs/1512.01360) [gr-qc] | [PDF](#)

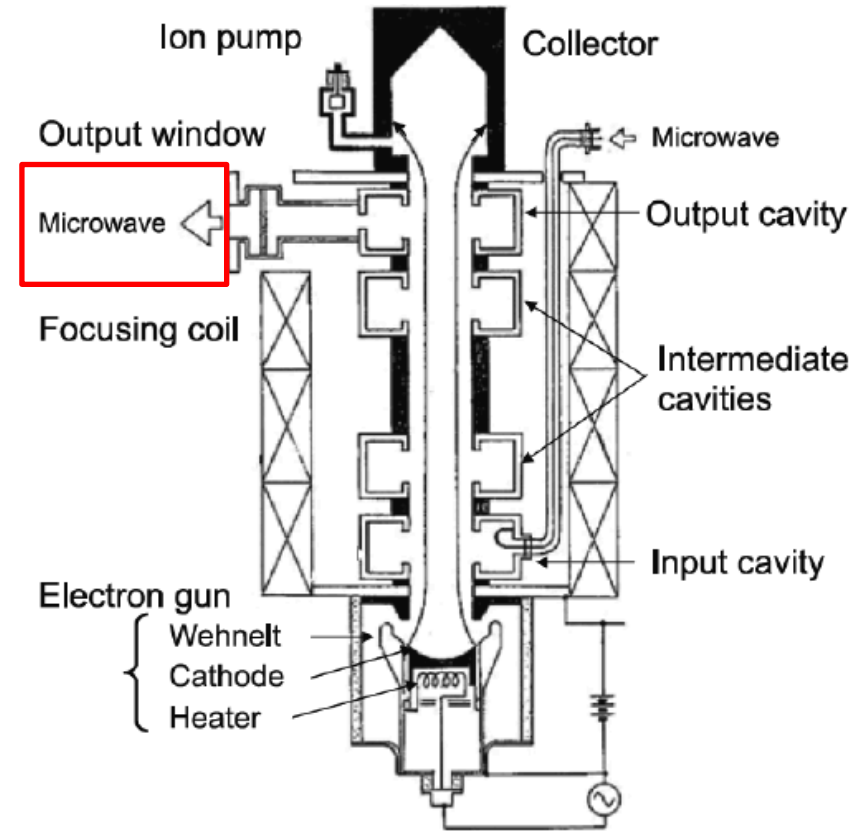
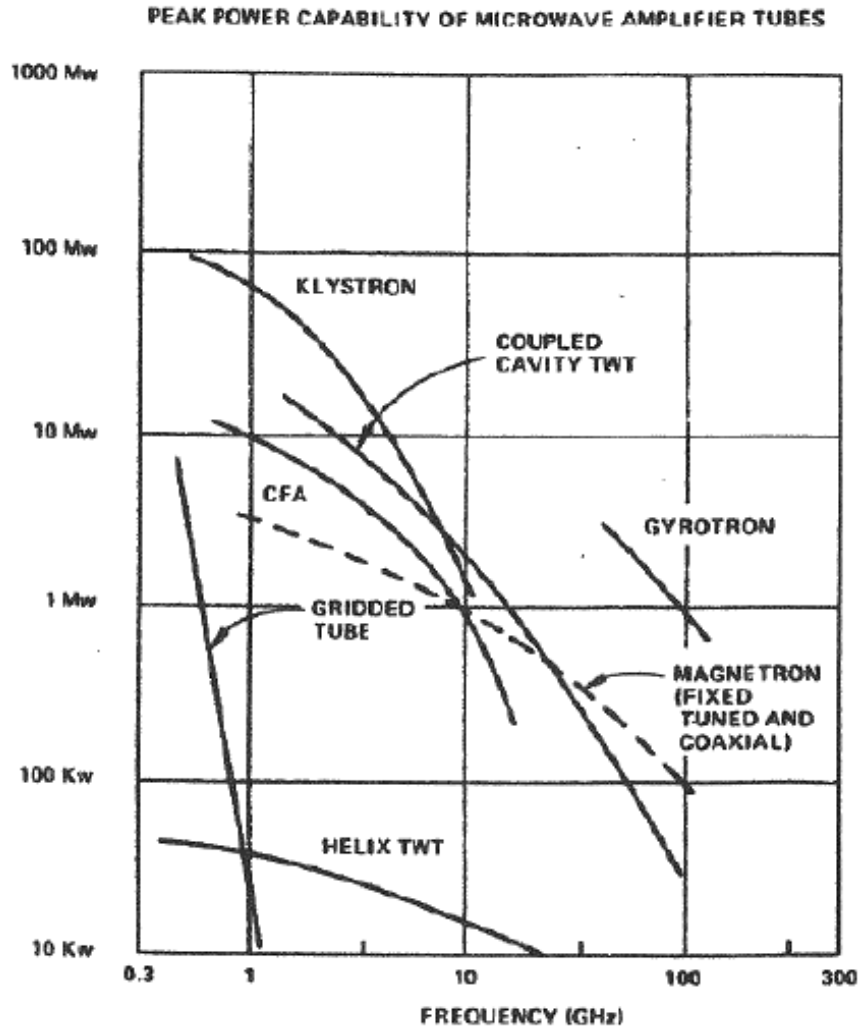


dilaton mass $m_\sigma = 0.15 \sim 0.59 \mu\text{eV} @ m_{\text{Higgs}} = 126 \text{ GeV}$

$$-L = \frac{1}{4} \frac{g}{M} F^{\mu\nu} F_{\mu\nu} \sigma$$

dilaton-photon coupling $g/M = (0.5 \sim 2) / (3\pi) (\alpha_{\text{qed}} / M_p)$

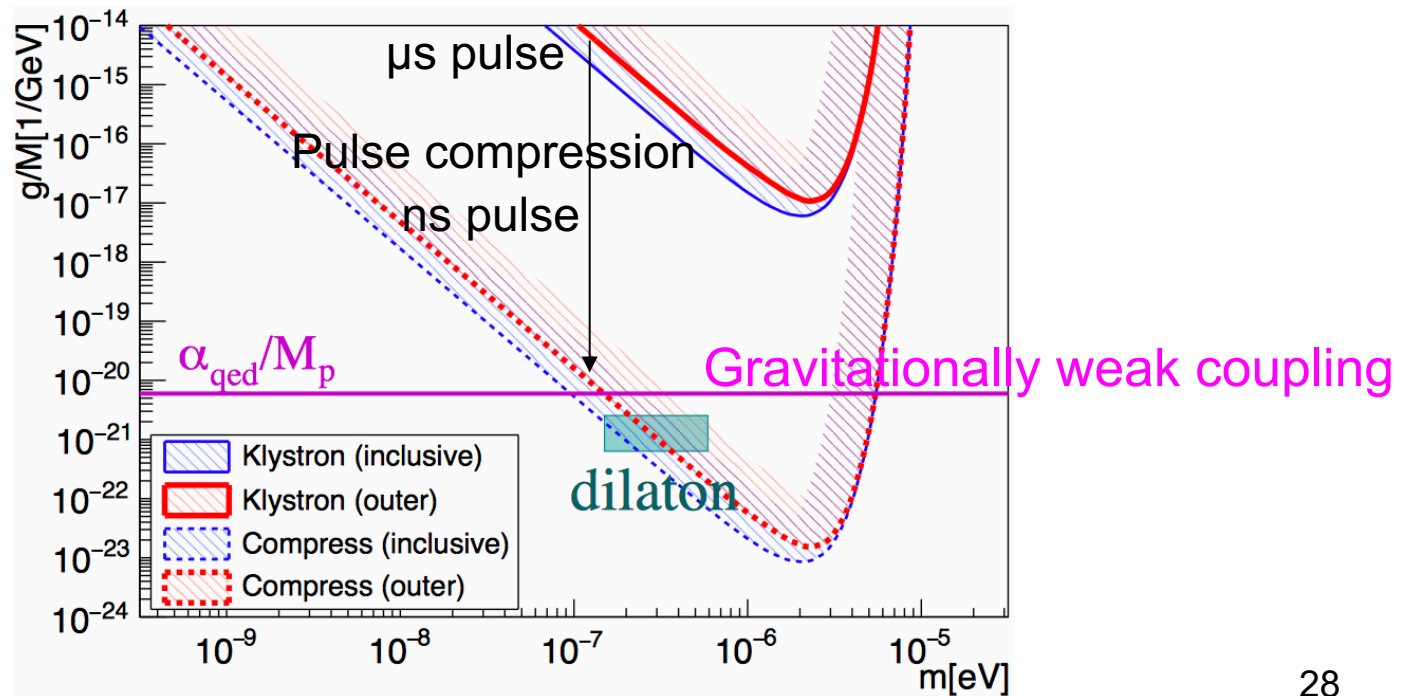
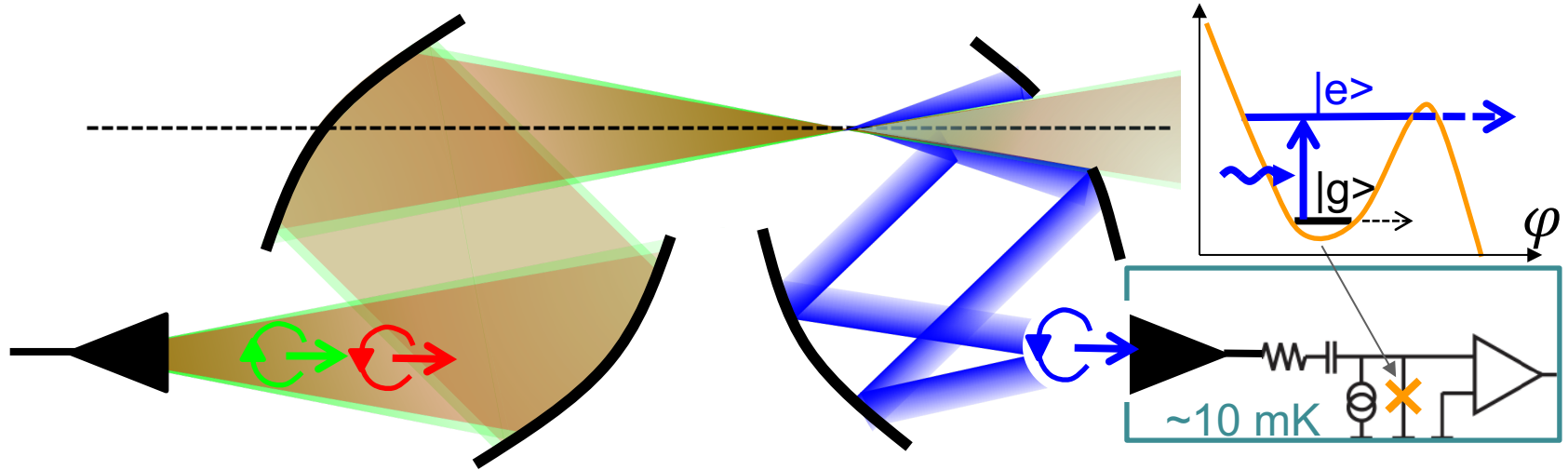
Klystron as a 10^{-5} eV photon source



J. Plasma Fusion Res. Vol.86 (2010)

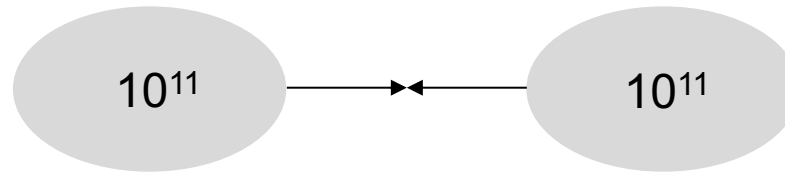
[3] Microwaves made simple: principles and applications/ the staff of the Microwave Training Institute, ed. By W. S. Cheung and F. H. Levin, Artech House, Boston 1985.

Focused radar collider



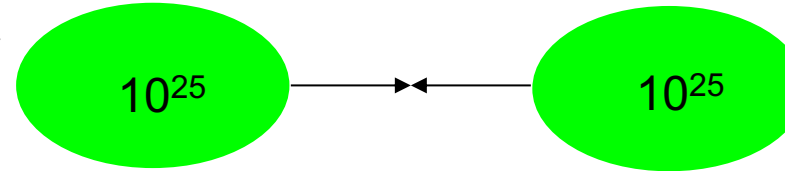
Comparison with charged particle colliders

Space-charge
limitation exits

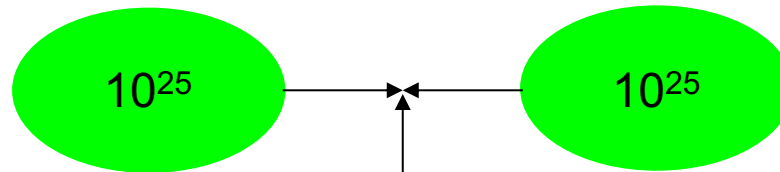


$$(10^{11})^2$$

No limit on # of photons*
 $10^{25} \sim 100\text{J}@\text{GHz}$



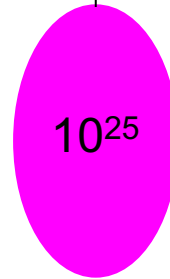
$$(10^{25})^2$$



Inducement

$$(10^{25})^3$$

* Threshold for real vacuum polarizations
far beyond the cutting-edge intensity



**Accessibility to
coupling even weaker
than that of gravity**

Summary

Stimulated resonant photon-photon colliders can extend the present horizon of particle physics and cosmology

huge gap
in coupling

| force | nuclear | strong | electro magnetic | weak | gravitational |
|---------------|-------------------|-------------------|------------------|-------------------|-------------------|
| strength | 10 | 0.1 | 1/137 | 10 ⁻⁵ | 10 ⁻³⁸ |
| distance (cm) | 10 ⁻¹³ | 10 ⁻¹³ | ∞ | 10 ⁻¹⁶ | ∞ |
| potential | exp(-mr) / r | a / r + b r | 1 / r | exp(-mr) / r | 1 / r |
| gauge boson | pion | gluon | photon | W / Z | graviton |
| theory | Yukawa | QCD | QED | Electroweak | Relativity |